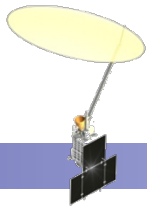


SMAP Radiometer L1/L2 Cal/Val Overview

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In collaboration with: P. Mohammed,
J.Peng, J.Piepmeier



L1 radiometer cal requirements are derived from....

- The **L2_SM_P** algorithm requires calibrated absolute Tb (L1B's output) as an input parameter. Also, the **L2_SM_A/P** algorithm fundamentally depends on the passive-only algorithm to retrieve SM (radar is only used for resolution enhancement--aka disaggregation).
- The **L3_SM_P** and **L3_SM_A/P** products are composites of the respective **L2 SM** products, so these **L3** products have the same requirement for calibrated absolute Tb as the corresponding **L2** products.
- The **L1C_TB** product is a gridded version of the L1B product, so L1C requires calibrated absolute Tb, too.
- Relevant existing req's: L2-SR-45, L2-SR-388, & others
- **So the following products depend on calibrated absolute Tb's:**
 - **L1B_TB, L1C_TB**
 - **L2_SM_P, L2_SM_A/P**
 - **L3_SM_P, L3_SM_A/P**

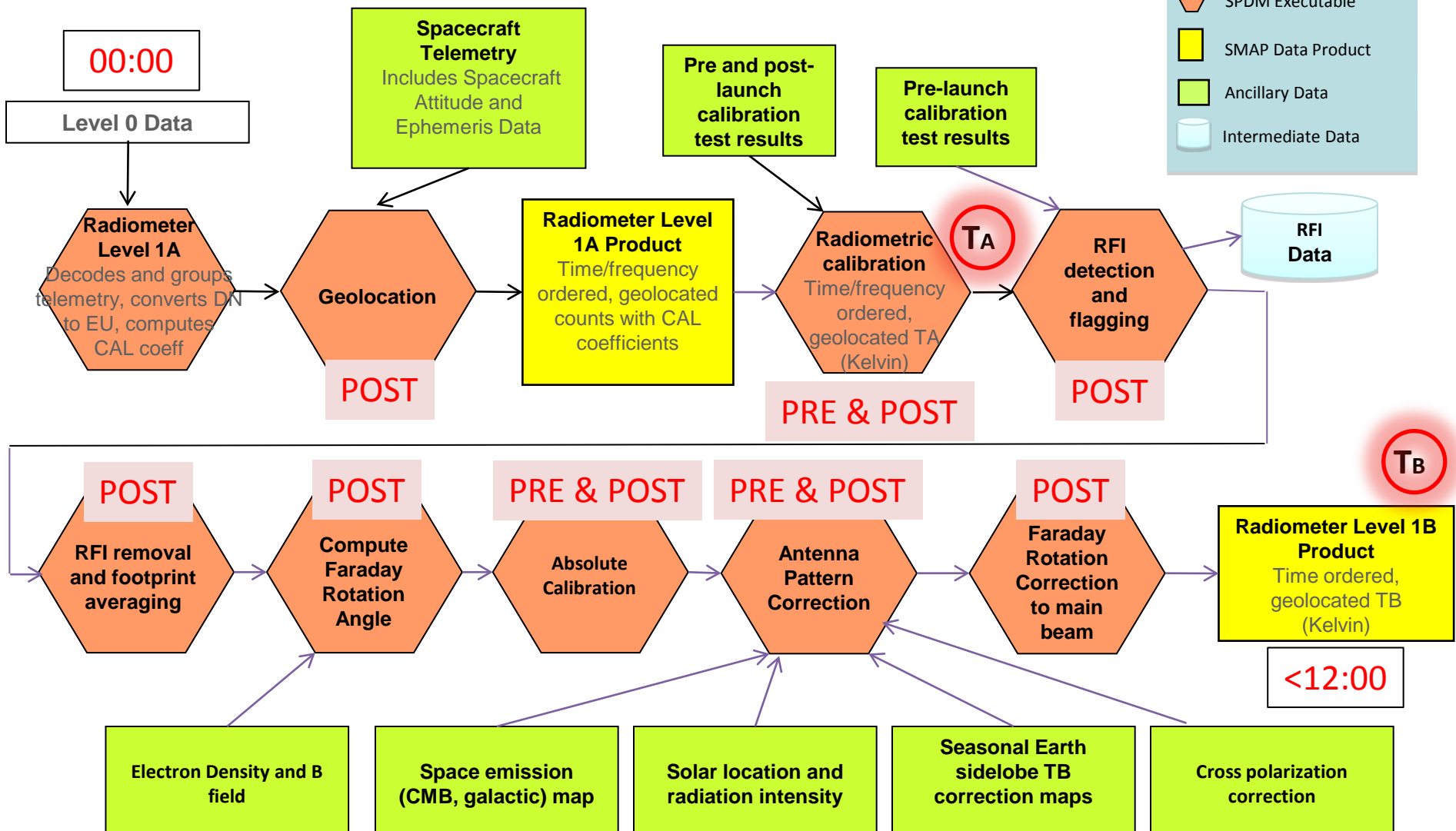
Radiometer Level 1A/B Processing Flow



Radiometer L1 Algorithm is driven by processing flow L0→L1A→L1B
 Radiometer Calibration is driven by information source (Pre vs. Post)

Key

- Telemetry and Input Data
- SPDM Executable
- SMAP Data Product
- Ancillary Data
- Intermediate Data

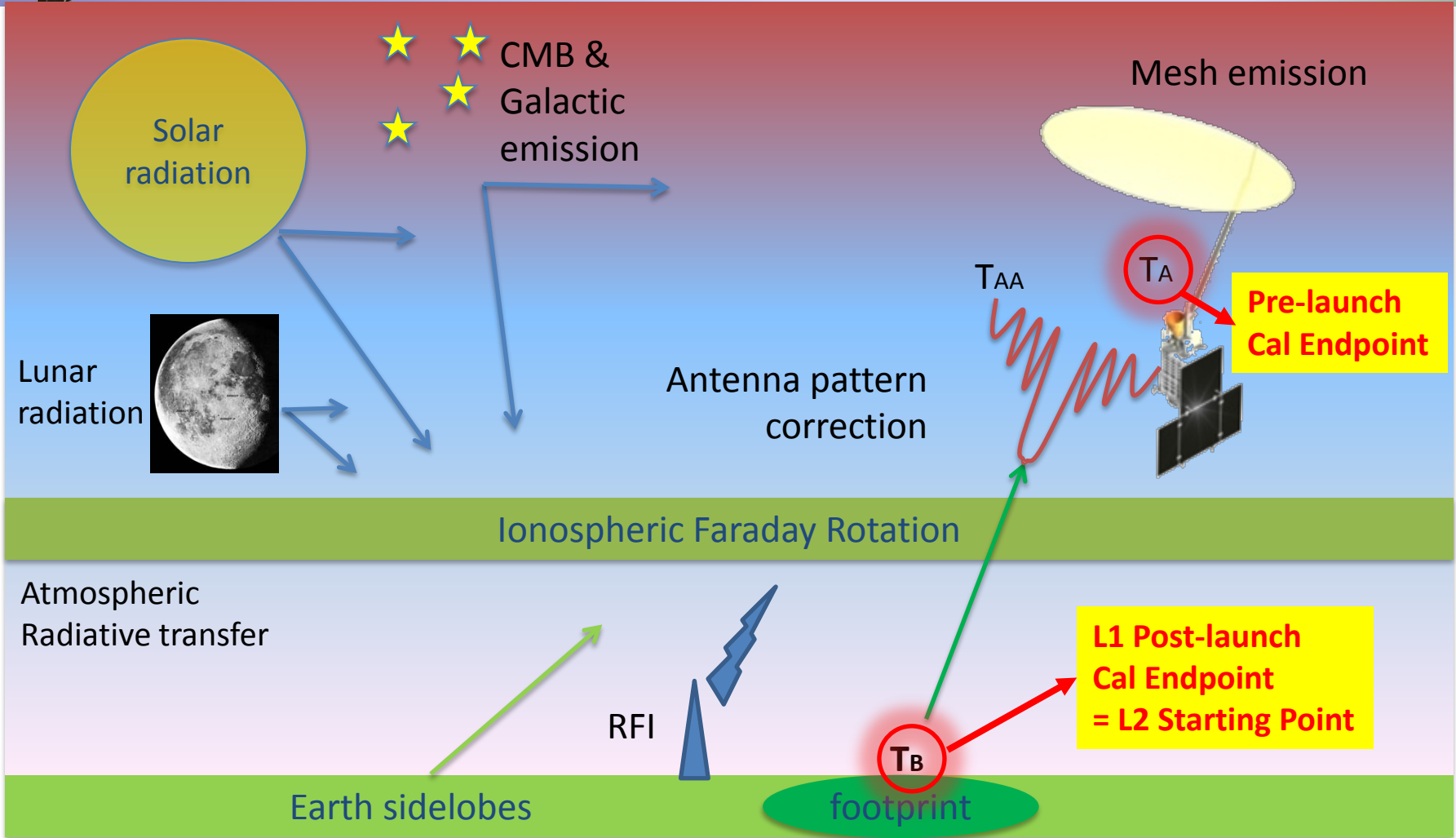


SMAP Radiometer L1 cal start & end points

L1 Post-launch cal starts where L1 Pre-launch calibration ends: radiometer calibrated up through the feedhorn.

L1 Post-launch calibration ends with Tb calibrated & corrected to bottom of atmosphere—i.e. where L2 algorithm begins.

SMAP Radiometer L1 Post-Launch Calibration Effects Considered

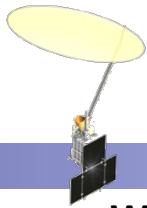




SMAP Radiometer L1 Cal/Val-related Top-Level Requirements



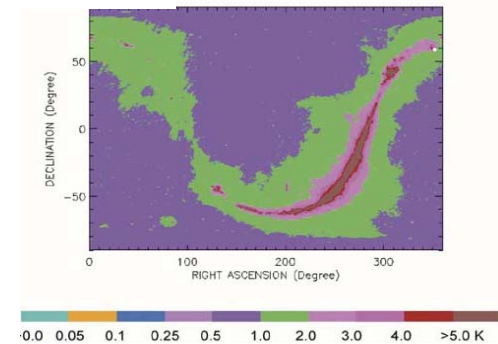
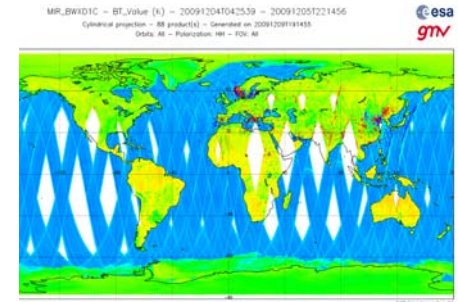
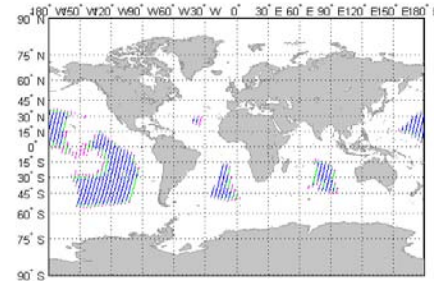
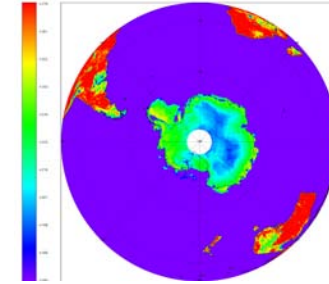
- L2-SR-45: Radiometer Grid Measurements shall have an uncertainty (including instrument precision, post-launch calibration error, and RFI-induced error) of 1.3 K or less (1-sigma) in the H and V channels.
 - Spatial & temporal = global, 6-month avg (clarified by SDT in Feb, 2010)
- L2-SR-34: The SMAP radiometer shall measure H, V, and 3rd and 4th Stokes parameter brightness temperatures. [T3 for Faraday Rot., T4 for RFI]
- L2-SR-295: Radiometer Level 1B Processing shall include compensation for effects of antenna sidelobes (outside the radiometer antenna main beam), cross-polarization, Faraday rotation, atmospheric effects (excluding rain), and solar, galactic and cosmic radiation.
- L2-SR-388: The SMAP project shall conduct a calibration and validation (Cal/Val) program: (a) Pre-launch, to ensure the development of robust science algorithms, and (b) Post-launch, to verify and characterize the accuracies of the delivered science data products



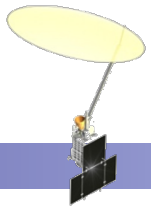
Absolute Tb calibration



- **Why?** To meet needs of multiple SMAP data products
- **Methods**
 - External targets since an on-board beamfilling target is not feasible
 - Use at least 2 views of external targets w/different Tb's (2-point cal)
- **Candidate targets**
 - **Antarctica** (~200K)
 - SMOS work shows promise
 - Near pole → Visible multiple times/day
 - No maneuver or special mode required
 - RFI unlikely
 - **Ocean** (~100K)
 - leverage Aquarius work on same
 - No maneuver or special mode required
 - Absolute Tb cal depends on ocean model accuracy
 - RFI still possible
 - **cold space** (~10K)
 - Requires s/c maneuver; limits how often (monthly)
 - Avoid looking toward celestial source regions
 - RFI unlikely



Faraday rotation correction



$$\Omega_f = 2.62 \times 10^{-13} \lambda^2 \int n_e B_{\parallel} ds$$

- **Objective**

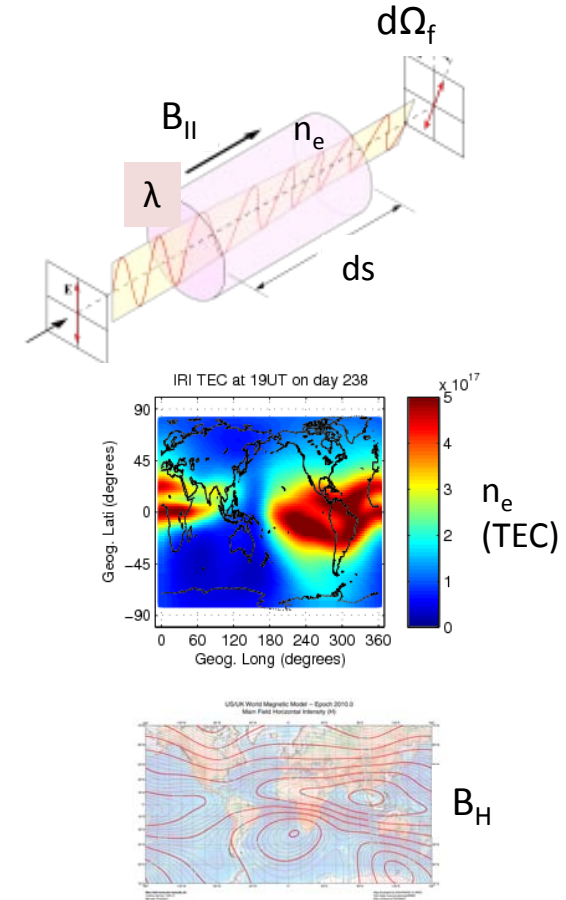
- Correct for polarization rotation due to ionosphere
- relevant to L2-SR-295, L2-SR-45

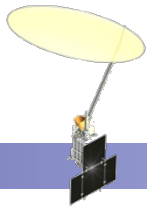
- **Method**

- Baseline is same as Aquarius, adapted for SMAP conical scan; leverage Hydros L1B work; check using SMOS
- n_e from GSFC using International Reference Ionosphere (IRI)
- B-field from International Geomagnetic Reference Field (IGRF)
- Compute Faraday rotation & rotate SMAP polarization basis

- **Expected outputs**

- Corrections to Tb's for APC and absolute Tb cal



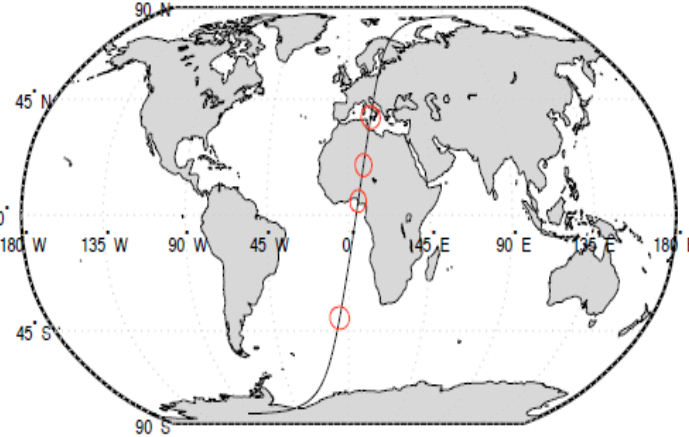


SMAP T_B Forward Simulator

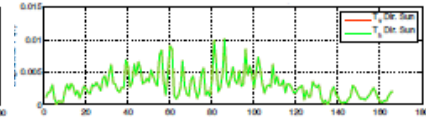
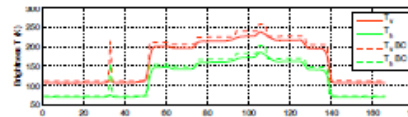


(used heavily by Antenna Pattern Correction)

- Strategy: leverage Aquarius simulator as much as possible
- Major differences vs. Aquarius
 - Conical scanning vs. fixed beam
 - Different antenna pattern
 - Land focus vs. ocean
- Proof-of-concept exercise completed
 - used sources already in Aquarius simulator
 - Simulated 4 cases along 1 descending orbit
 - Examined results for L1 error budget implications

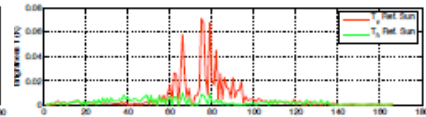
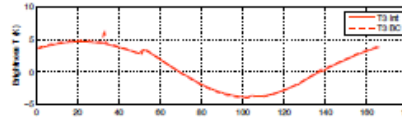


TV, TH



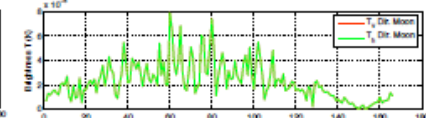
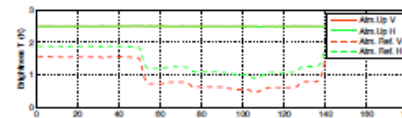
Solar direct

T3



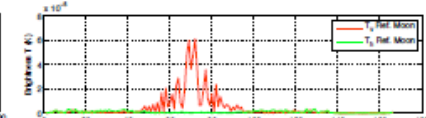
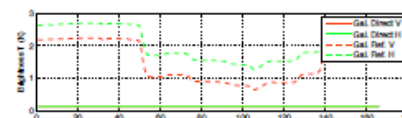
Solar reflected

Atmos



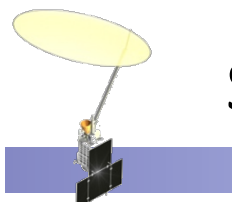
lunar direct

galactic

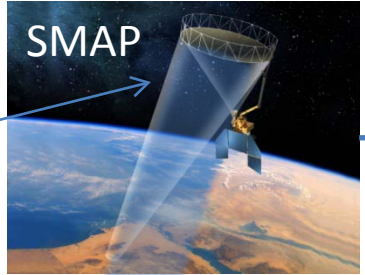


lunar reflected

SMAP radiometer L1/L2 integrated cal/val flow



Unwanted Contributions (sun, moon, sidelobes, etc)



L0 Radiometer counts

L1 Calibration Process aka "L1 Tb algorithm"

L1 passive val consists of comparing **these** and then adjusting **this**.

L1 Brightness Temperatures (Tb)

True Brightness Temperatures (fcn of temp, moisture, EIA, etc)



L1 TB Validation: Measure TB

Upscale to SMAP pixel size

TB "truth" at SMAP resolutions

L2 passive algorithm

SMAP L2 passive retrieved SM



L2 SM Validation: Measure SM

Measured SM from cal/val field experiment

Upscale to SMAP pixel size

SM "truth" at SMAP resolutions

L1/L2 passive Cal/val consists of comparing **these** and then adjusting **these**.